

# RICCI and Matter Collineations of SOM-ROY Chaudhary Symmetric Space Time

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## ABSTRACT

This paper is devoted to explore the RICCI and MCs (Matter Collineations of the Som-Ray Chaudhary spacetime). The spacetime under consideration is one of the spatially homogeneous and rotating spacetimes. Collineations are the some kinds of the Lie symmetries. To discuss the required collineations we have used the RICCI and energy momentum tensors. As the RICCI tensor is formulated from the metric tensor, it must possess its symmetries. RCs (RICCI Collineations) leads to conservation laws. On the other hand for the distribution of matter in the spacetimes, the symmetries of energy momentum tensor or MCs provides conservation laws on matter field. Throughout this paper, these collineations are discussed by vanishing Lie derivative of RICCI and energy momentum tensors respectively. Complete solution of the RCs and MCs equations, which are formed in the result of vanishing Lie derivative are explored. Studying all these collineations in the said spacetime, it has been shown that RCs of the spacetime form an infinite dimensional vector space where as MCs are Killing vector fields.

**Key Words:** RICCI Collineations, Matter Collineations, Killing Vector Fields, Infinite Dimensional Vector Space.

## 1. INTRODUCTION

Collineations are the geometric assumptions which concern the “symmetries” of the metric defined by the Lie symmetries of either the metric itself or the tensors defined from the metric [1,2].

Theory of general relativity is basically the theory of gravitation and is nonlinear due to curved spacetimes. Einstein expressed this phenomenon in terms of set of gravitational field equations known as EFEs (Einstein Field Equations) are:

$$R_{ab} - \frac{1}{2}Rg_{ab} + \Lambda g_{ab} + \kappa T_{ab}$$

where  $a, b, = 0, 1, 2, 3$ .

The left hand side of EFEs describes the geometry while the RHS is totally related to physics. In these equations,  $\Lambda$  is called cosmological constant and  $\kappa = 8\pi G/C^4$  is called coupling constant [3]. Due to high non linearity, it is very difficult to find the exact solutions of the system. To overcome this difficulty we introduce symmetry

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restrictions. These symmetry restrictions for mechanical problems reduce the degrees of freedom. Reduction of degrees of freedom simplifies the problem. The important symmetries which are frequently used in general relativity are KVF's (Killing Vector Fields), HVF's (Homothetic Vector Fields), CFV's (Conformal Vector Fields), CC's (Curvature Collineations), RC's and MC's [4]. Therefore, it is necessary to discuss the symmetries of different kinds in general relativity. Most of the symmetries give the conservation laws. Owing to that in this paper we have discussed two types of the important symmetries, namely RC's and MC's of said spacetime. It is obvious from the EFE's that RICCI and matter collineations are necessary to find the exact solutions of EFE's. MC's are totally related to the physical properties of the spacetimes, whereas RC's are concerned with the geometry of spacetimes [2,5]. There are many approaches to discuss RC's and MC's but the technique used to find RC's and MC's is vanishing Lie derivative of RICCI and energy momentum tensors along with the specified vector fields [6]. In this work, the system of nonlinear coupled partial differential equations generated by taking Lie derivative is solved by direct integration technique [5,7]. Normally in general relativity, Mis used for the manifold, which is four dimensional, connected, Hausdorff spacetime. The metric  $g$  which contains all of the information about the geometry of the manifold is taken as Lorentzian with signature  $(-, +, +, +)$ . The RICCI and energy momentum tensors in component form which are connected with metric tensor  $g_{ab}$ , are denoted by  $R_{ab}$  and  $T_{ab}$  respectively. Comma, semicolon and  $L$  are the symbols used for partial, covariant and Lie derivatives respectively [5].

## 2. RESULTS FOR RICCI COLLINEATIONS

Consider a Som-Ray Chaudhary spacetime in the curvilinear coordinates  $(t,r,\theta,z)$  labeled by  $(x^0, x^1, x^2, x^3)$  respectively. Here, superscripts are not the powers. The line element is [8]:

$$ds^2 = - dt^2 + dr^2 + r^2(1-r^2) d\theta^2 + dz^2 + 2r^2 dt d\theta \quad (1)$$

The RICCI tensor and RICCI scalar in the usual notations are defined as  $R_{ab} = R^c_{acb}$ ,  $R = g^{ab} R_{ab}$ . The non-zero components of RICCI tensor and the value of RICCI scalar are given by  $R_{00} = 2 = R_{11}$ ,  $R_{02} = R_{20} = 2r^2$ ,  $R_{22} = 2r^4 + 2r^2$ ,  $R = 2$ . The vector field  $\chi$  is said to be RC's if it satisfies the relation [6].

$$L\chi R_{ab} = 0 \quad (2)$$

Equivalently,

$$R_{ab,c} \chi^c + R_{bc} \chi^c_a + R_{ac} \chi^c_b = 0$$

where  $a, b, c = 0, 1, 2, 3$ .

The explicit form of the Equation (2) can be written as a set of nonlinear coupled partial differential equations as:

$$\chi^0_{,0} - r^2 \chi^2_{,0} = 0 \quad (3)$$

$$\chi^0_{,1} + \chi^1_{,0} - r^2 \chi^2_{,1} = 0 \quad (4)$$

$$2r\chi^1 - \chi^0_{,2} + r^2 \chi^0_{,0} - (r^4 + r^2) \chi^2_{,0} + r^2 \chi^2_{,2} = 0 \quad (5)$$

$$\chi^0_{,2} - r^2 \chi^2_{,2} = 0 \quad (6)$$

$$\chi^1_{,1} = 0 \quad (7)$$

$$r^2 \chi^0_{,1} - \chi^1_{,2} - (r^4 + r^2) \chi^2_{,1} = 0 \quad (8)$$

$$\chi^1_{,3} = 0 \quad (9)$$

$$(2r^2 + 1) \chi^1 - r \chi^0_{,2} + (r^3 + r) \chi^2_{,2} = 0 \quad (10)$$

$$\chi^0_{,3} - (r^2 + 1) \chi^2_{,3} = 0 \quad (11)$$

From Equations (3,4,7,9), we get

$$\left. \begin{aligned} \chi^0 &= -\frac{1}{2} r A_t^2(t, \theta) + \int r^2 A_r^3(r, \theta, z) dr + A^5(\theta, z) \\ \chi^1 &= A^2(t, \theta) \\ \chi^2 &= -\frac{1}{2} A_t^2(t, \theta) + A^3(r, \theta, z) \end{aligned} \right\} \quad (12)$$

Where  $A^2(t, \theta)$ ,  $A^3(r, \theta, z)$  and  $A^4(\theta, z)$  are functions of integration. To avoid from the lengthy calculations we

only present the results. So, if one proceeds further, using the information obtained from Equation (12) in the remaining equations, we have the result.

$$\left. \begin{aligned} \chi^0 &= r(c_5 \cos \theta + c_6 \sin \theta) + c_8 \\ \chi^1 &= -c_5 \sin \theta + c_6 \cos \theta \\ \chi^2 &= -\frac{1}{r}(c_5 \cos \theta + c_6 \sin \theta) + c_4 \end{aligned} \right\} \quad (13)$$

Finally, the RCs become

$$\left. \begin{aligned} \chi^0 &= r(c_5 \cos \theta + c_6 \sin \theta) + c_8 \\ \chi^1 &= -c_5 \sin \theta + c_6 \cos \theta \\ \chi^2 &= -\frac{1}{r}(c_5 \cos \theta + c_6 \sin \theta) + c_4 \\ \chi^3 &= \xi(x^a), a = 0,1,2,3 \end{aligned} \right\} \quad (14)$$

Here  $\xi(x^a)$  is an arbitrary function and  $C_4, C_5, C_6, C_8$  are arbitrary constants. RCs after subtracting the Killing vector fields are:

$$\chi = (0,0,0,\xi(x^a)), a = 0,1,2,3 \quad (15)$$

Clearly RCs due to existence of the arbitrary function form an infinite dimensional vector space.

### 3. MAIN RESULTS FOR MATTER COLLINEATIONS

The expression of energy momentum tensor by considering the cosmological constant equals zero and the coupling constant equal to unity, takes the form [7]:

$$T_{ab} = R_{ab} - \frac{1}{2} R g_{ab} \quad (16)$$

The non-zero components of energy momentum tensor are:

$$T_{00} = 3, T_{11} = 1, T_{02} = T_{20} = -3r^2, T_{22} = 3r^4 + r^2, T_{33} = -1 \quad (17)$$

The vector field  $\eta$  is said to be MCs if the Lie derivative along this vector field vanishes. Mathematically it is represented by the relation.

$$\frac{LT_{ab}}{\eta} = 0 \quad (18)$$

Expansion of the Equation (18) is given by:

$$T_{ab,c} \eta^c + T_{bc} \eta_{,a}^c + T_{ac} \eta_{,b}^c = 0$$

Here, a, b, c = 0, 1, 2, 3 and comma shows the partial derivative. The explicit form of the above Equation (18) with the help of Equation (17) can be written as set of ten nonlinear coupled partial differential equations:

$$\eta_{,0}^0 - r^2 \eta_{,0}^2 = 0 \quad (19)$$

$$3\eta_{,1}^0 + \eta_{,0}^1 - 3r^2 \eta_{,1}^2 = 0 \quad (20)$$

$$6r\eta^1 - 3\eta_{,2}^0 + 3r^2 \eta_{,0}^0 - (3r^4 + r^2) \eta_{,0}^2 + 3r^2 \eta_{,2}^2 = 0 \quad (21)$$

$$3\eta_{,0}^3 - 3\eta_{,3}^0 + 3r^2 \eta_{,3}^2 = 0 \quad (22)$$

$$\eta_{,1}^1 = 0 \quad (23)$$

$$3r^2 \eta_{,1}^0 - \eta_{,2}^1 - (3r^4 + r^2) \eta_{,1}^2 = 0 \quad (24)$$

$$3\eta_{,3}^1 - \eta_{,1}^3 = 0 \quad (25)$$

$$(6r + r) \eta^1 - 3r^2 \eta_{,2}^0 + (3r^4 + r^2) \eta_{,2}^2 = 0 \quad (26)$$

$$3\eta_{,0}^3 + 3r^2 \eta_{,3}^0 - (3r^4 + r^2) \eta_{,3}^2 = 0 \quad (27)$$

$$\eta_{,3}^3 = 0 \quad (28)$$

From Equations (19,20,23) and Equation (28), we get:

$$\left. \begin{aligned} \eta^0 &= -\frac{1}{6} z r A_r^3(t, \theta) + \int r A^9(t, \theta) dt + \int r^2 A_r^7(r, \theta, z) dr \\ &\quad - \frac{r}{3} A^{10}(\theta) + A^{12}(\theta, z) \\ \eta^1 &= z A^3(t, \theta) - 6 \int \{A^9(t, \theta) dt\} dt + t A^{10}(\theta) + A^{11}(\theta) \\ \eta^2 &= -\frac{1}{6r} z A_r^3(t, \theta) + \int \frac{1}{r} A^9(t, \theta) dt + A^7(r, \theta, z) \\ \eta^3 &= r A^3(t, \theta) + A^4(t, \theta) \end{aligned} \right\} \quad (29)$$

Here all  $A$ 's are the functions of integration to be determined with the help of direct integration technique. Again to avoid from tedious and lengthy calculation, only results are presented [6]. Therefore, the solution of the system of equations from Equations (19-28) is given:

$$\left. \begin{aligned} \eta^0 &= r(c_{11} \cos \theta + c_{12} \sin \theta) + c_{14} \\ \eta^1 &= -c_{11} \sin \theta + c_{12} \cos \theta \\ \eta^2 &= -\frac{1}{r}(c_{11} \cos \theta + c_{12} \sin \theta) + c_{10} \\ \eta^3 &= c_6 \end{aligned} \right\} \quad (30)$$

Here all  $C, s$  are the arbitrary constants. Clearly, this system satisfies the Killing equation. Therefore, MCs of the given spacetime are the Killing vector fields.

#### 4. CONCLUSIONS

From the above discussion we come to the following conclusions:

- (i) When the RCs of the above spacetime has been discussed, it is found that RCs form an infinite dimensional vector space as mentioned in Equation (15).
- (ii) In case of studying MCs of the above spacetime, it has been found that MCs are precisely the Killing vector fields as shown in Equation (30).

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